

Unfrozen hyperscalars and Supersymmetry

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Work in collaboration with
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Hypermultiplets are rather simple:

$\left. \begin{array}{l} \text{hyperini} \\ \text{hyperscalars} \end{array} \right\} \begin{array}{l} 2 \\ 9 \end{array}$

The geometry describing the coupling between the hyperscalars is Quaternionic Kähler

$n_0 > 1$: $J^r J^s = -\delta^{rs} + \epsilon^{rst} J^t$

$$H(J^r X, J^s Y) = H(X, Y)$$

$$\nabla_x J^r = \underbrace{-2 \epsilon^{rst} A_x^s}_{\text{SU}(2)\text{-connection}} J^t$$

Implies \Rightarrow H is Einstein

\Rightarrow $\text{Hol} \subseteq \text{Sp}(1) \cdot \text{Sp}(n_0)$

if A is pure gauge, we recover hyperkähler.

$n_0 = 1$: A 4-dimensional manifold, with metric H , is said to be Quaternionic Kähler if H is

- Einstein & Selfdual

N.B. Selfduality means that the Weyl tensor is selfdual!

Coupling to Supergravity

- To the action only through a σ -model

$$S = \int \sqrt{g} [R + \dots + H_{xy} \partial q^x \partial q^y]$$

- the susy variations of the hypers:

$$\delta d^\alpha = U_x^{\alpha I} \not{\partial} q^x \epsilon_I$$

- the susy rule for the gravitino

$$\delta \psi_\mu = (\nabla_\mu + \not{\partial}_\mu q^x A_x) \epsilon + \dots$$

Effects: Following ideas by K.P. Tod, one assumes the existence of a solution to the KSE, and uses those to construct tensor-fields. These tensor fields obey 2 types of constraints

- 1) Non-differential. I.e. Fierz Identities
- 2) Differential Constraints: w.r.t. the case without hypers, this is where the misery starts: The $\not{\partial}(\epsilon)$ part in the gravitino variation turns up!

$N=1$ $d=5$ Supergravity:

Spinors ϵ_I ($I=1,2$) are Symplectic M.

$$\rightarrow f = i \bar{\epsilon}_I \epsilon^I$$

$$V^a = i \bar{\epsilon}_I \gamma^a \epsilon^I$$

$$\bar{\Phi}_{ab}^r = \sigma_I^r \bar{\epsilon}_J \gamma_{ab} \epsilon^I$$

And the Fierz identities imply amongst other things:

$$V_a V^a = f^2 \quad V^a \bar{\Phi}_{ab}^r = V^a r \bar{\Phi}_{abc}^r = 0$$

$$\bar{\Phi}_a^r \bar{\Phi}_{cb}^s = -\delta^{rs} (\eta_{ab} f^2 - V_a V_b) - \epsilon^{rst} f \bar{\Phi}_{ab}^t$$

As is usual $\nabla_{(a} V_{b)} = 0 \Rightarrow V^a \partial_a \equiv \partial_t$

$$ds^2 = f^2 (dt + \omega)^2 - f^{-1} h_{mn} dx^m dx^n$$

$\bar{\Phi}^r$ only has m, n indices and induces a quaternionic structure on h_{mn}

Two very important differential constraints are

$$\textcircled{\text{I}} \quad \partial_m q^x = \sum_r \Phi_m^r \partial_n q^r J_r^x$$

$$\textcircled{\text{II}} \quad \nabla_m \Phi_{np}^r = -2 \epsilon^{rst} A_m^s \Phi_{np}^t$$

$\textcircled{\text{I}}$ States that q is a quaternionic map. One can see that $\textcircled{\text{I}}$ implies that q is also harmonic, which is also the condition imposed by its equation of motion:

Before I forget: In the timelike case all the eq. of motion are satisfied for a configuration that solves the KSEs.

$\textcircled{\text{II}}$ States that part of the spin connection of h_{mn} is related to the pull back of the $su(2)$ connection on the hypervariety

Oh good! The problem is intertwined

There are solutions, though.

Variations on a Solution by Joung, Kaya & Sergin hep-th/0608034

Consider the metric

$$ds^2 = \frac{dq^x dq^x}{(1 + \lambda q^2)^2}$$

$$x = 1, \dots, 4$$

$$q^2 = \sum_x q^x q^x$$

$$\lambda = +1 \quad S^4 = SO(5)/SO(4)$$

$$\lambda = -1 \quad H_4 = SO(1,4)/SO(4)$$

Case treated
by JKS

Both of these are quaternionic Kähler.

Needed later: $A^r = -\lambda \frac{q^x}{1 + \lambda q^2} J_{x4}^r dq^4$

where J_{x4}^r is the quaternionic structure on flat space.

Take as the metric on the Base space another conformally flat one

$$h_{mn} dx^m dx^n = \Omega^2(x) dx^m dx^m$$

and take the quaternionic structure to be the same as on S^4/H_4 .

then $q^m(x) = \frac{x^m}{x^4}$

$$[x^4 = (x^m x^m)^2]$$

Ω is fixed by the matching of connections to be

$$\Omega = (1 + dq^2)^{1/3} = \left(1 + \frac{\lambda}{r^6}\right)^{1/3}$$

$$\rightarrow ds_h^2 = \left(1 + \frac{\lambda}{r^6}\right)^{2/3} [dr^2 + r^2 dS^3]$$

Geometry:

$d=1$: The possible singularity at $r=0$ is illusory: all curvature invariants are Regular everywhere and the limiting geometry is but \mathbb{R}^4 .

$d=-1$: There is a true unprotected curvature singularity at $r=1$

So, what about solutions to SUGRA...
yes, I think I'd better say something..

$N=1$ $d=5$ Supergravity - no vector multiplet
- hypers on S^4/Ku .

A supersymmetric solution in this

setting is given by (more general ones are possible)

$$ds^2 = H^{-2} (dt + \omega)^2 + H \left(1 + \frac{\lambda}{r^6}\right)^{2/3} (dr^2 + r^2 dS^3)$$

$$A = -\sqrt{3} K^{-1} dt$$

$$q^m = \frac{x^m}{x^4} = \frac{x^m}{r^4}$$

ω is restricted by $f d\omega = - * [f d\omega] \Rightarrow \omega = 0$

H has to be harmonic w.r.t. h_{mn} .

$$H = 1 + \mathcal{O} \left(\frac{{}_2F_1\left(\frac{1}{3}, \frac{2}{3}; \frac{4}{3}; -\lambda r^{-6}\right)}{r^2} \right)$$

${}_2F_1$ is a Gauss hypergeometric function..

Some properties:

$\Rightarrow {}_2F_1(a, b; c; 0)$ is a constant

$$\Rightarrow \lim_{r \rightarrow \infty} H = 1$$

$\Rightarrow \frac{{}_2F_1\left(\frac{1}{3}, \frac{2}{3}; \frac{4}{3}; -\lambda r^{-6}\right)}{r^2}$ is monotonically decreasing.

What about $r=0, 1$??

First $d=-1$: then we are interested in

$${}_2F_1\left(\frac{1}{3}, \frac{2}{3}; \frac{4}{3}; 1\right) = \frac{\Gamma\left(\frac{1}{3}\right)\Gamma\left(\frac{4}{3}\right)}{\Gamma\left(\frac{2}{3}\right)}$$

$$\approx 1.76664$$

Nice, but remember the curvature singularity.

$d=1$: The calculation is a bit harder but the surprising result is

$$\lim_{r \rightarrow 0} \frac{{}_2F_1\left(\frac{1}{3}, \frac{2}{3}; \frac{4}{3}; -r^{-6}\right)}{r^2} = \frac{\Gamma\left(\frac{1}{3}\right)\Gamma\left(\frac{4}{3}\right)}{\Gamma\left(\frac{2}{3}\right)}$$

So if $d=1$ and $\rho > 0$, which is needed for positive mass, then our solution is Completely Regular.

A Quick Overview

Timelike

Null

$d=6$

$e^+e^- + B_4$

$d=5$

$T^2 + B_4$

\longleftrightarrow

$e^+e^- +$

B_3

$d=4$

$T^2 + B_3$

$\xrightarrow{\text{C-map}}$

$e^+e^- +$

B_2

